

#### Stars, Companions, and their Interactions A Memorial to Robert H. Koch

August 10-12, 2011 Villanova, PA USA

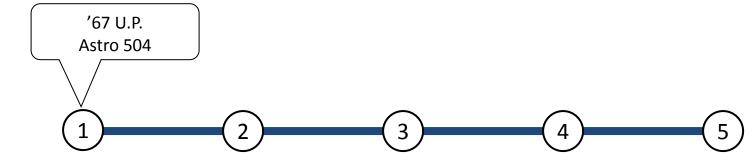
# **The Pennsylvania Polarimeters**

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## Abstract

- This report describes the inception, development and extensive use (over 30 years) of elliptical polarimeters at the University of Pennsylvania.
- The late 60's Mark I polarimeter design utilized oriented quarter wave plates and a calcite Foster-Clarke prism as the analyzer.
- Successive generations of automation and improvements included the late-70's optical redesign to utilize a photoelastic modulated wave plate and an Ithaco lock-in amplifier – the PEMP.
- The final design in 2000 concluded with a remotely operable device.
- Extensive studies of close binaries, pulsating hot stars, and luminous late-type variables.



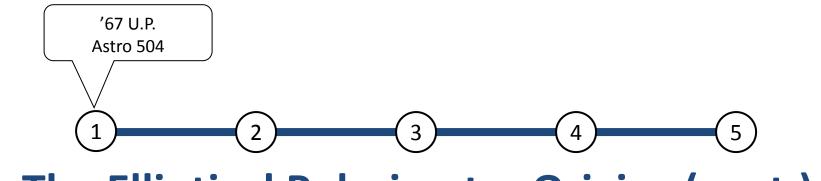


#### **1. The Elliptical Polarimeter Origins**

- The polarimeter began in 1967 as a termpaper design by George Wolf for Bill Blitzstein's AST 504 Astronomical Instrumentation class.
- Blitzstein then spoke with Brad Wood about the possibility of building the instrument.



**Bill Blitzstein** 



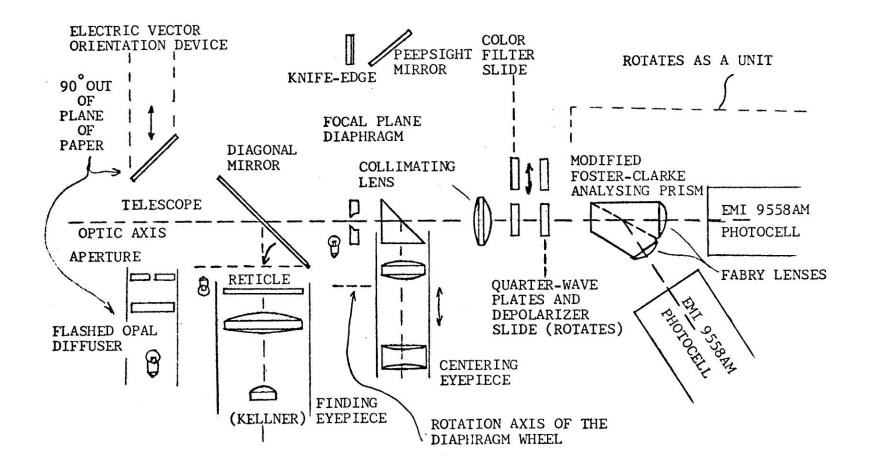
#### 1. The Elliptical Polarimeter Origins (cont.)

- Wood offered Wolf the remaining funds in an expiring NSF grant to purchase all of the optics.
- All optics were quickly purchased. Then the rest of the mechanical and electronic design was put together by Blitzstein and Wolf.
- In 1968-69 Bill Barrie did the machine work, and Bob Smith did the electronics to complete the instrument construction.

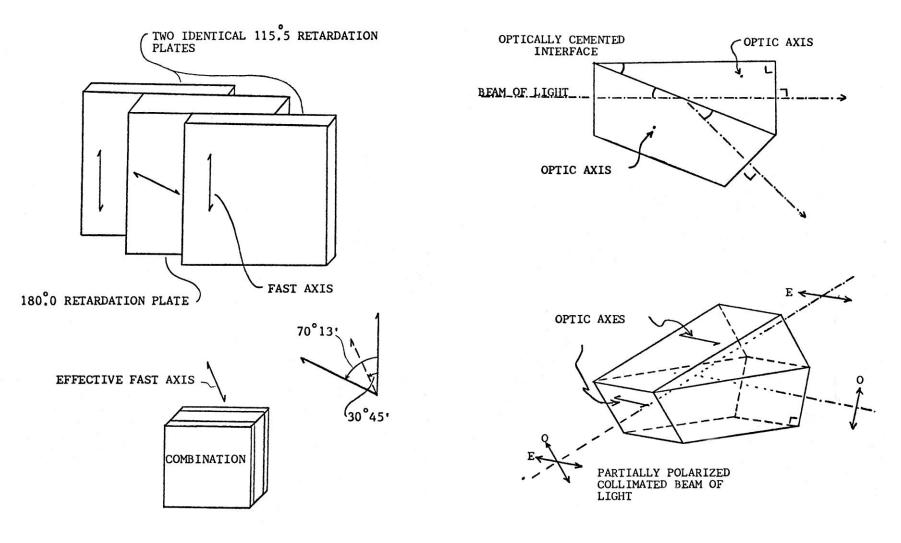


F. B. Wood

## Optical Components of the Mark I Polarimeter



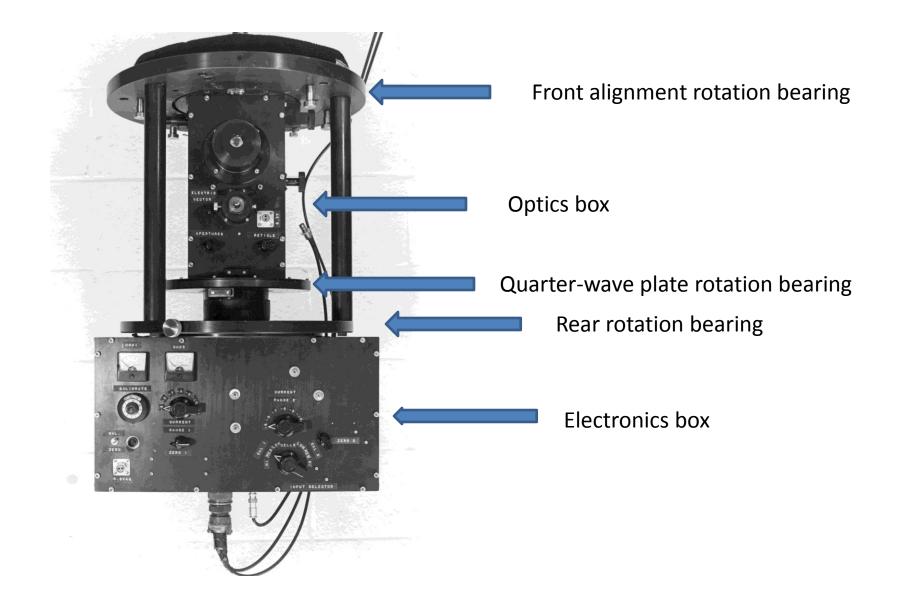
#### **Special Optical Component Details**



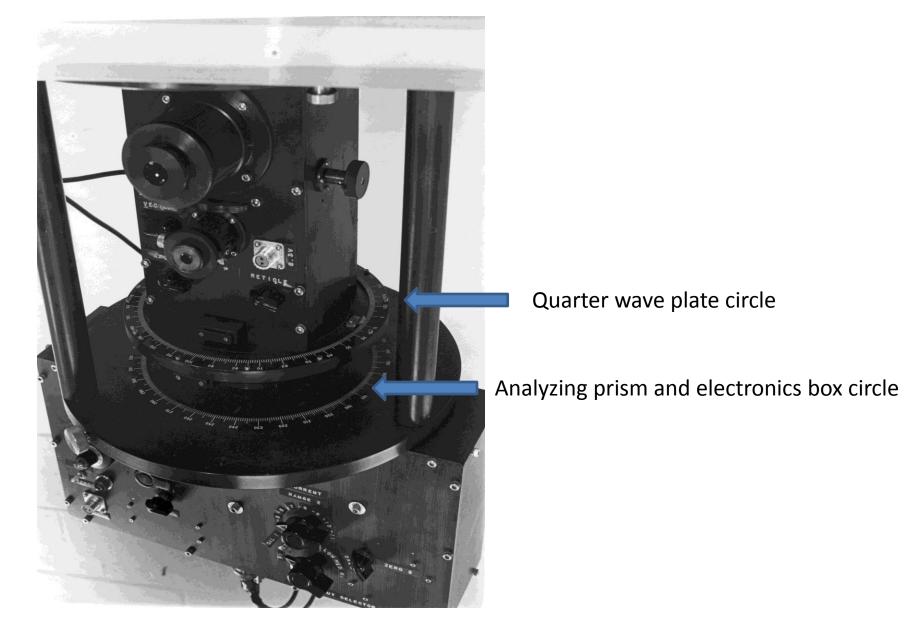
Achromatic Quarter-wave Plates

Calcite Foster-Clarke Prism

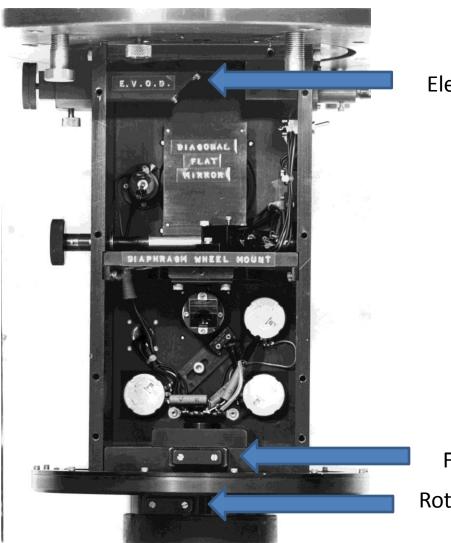
#### **The Completed Elliptical Polarimeter - 1969**



#### **Polarimeter Setting Circles**



#### **Optics box - inside**

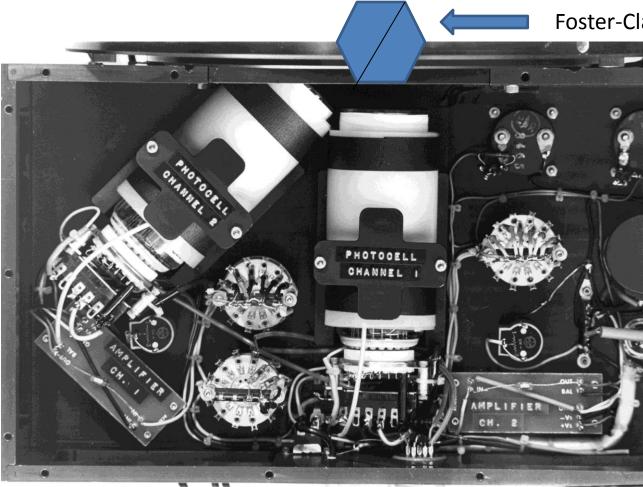


Electric –vector orientation device

Filter slide

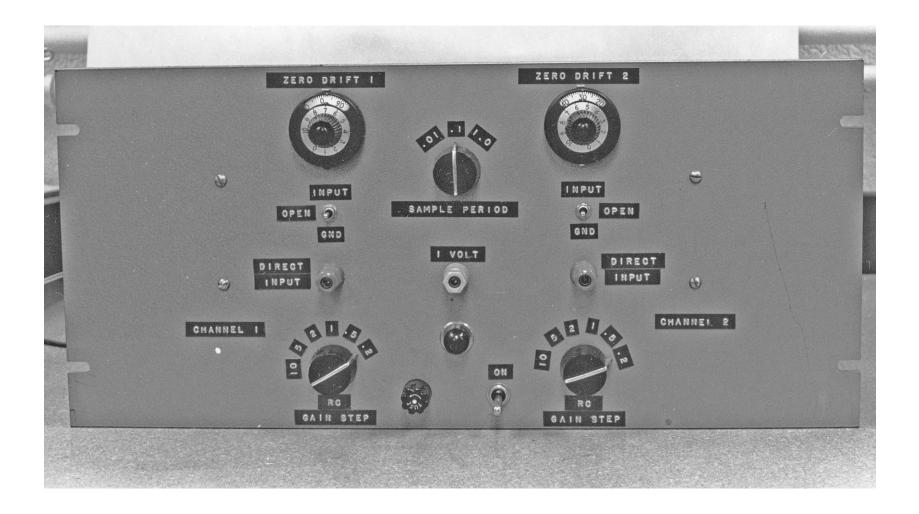
Rotating quarter-wave plate slide

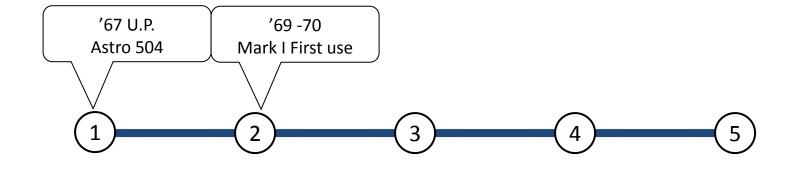
#### The Electronics Box - inside



Foster-Clarke Analyzing Prism

#### The Integrator – Power Supply – Relay Logic Unit





#### 2. Initial Use of the Mark I Polarimeter

- First use of the polarimeter was supposed to be a Southern Sky elliptical-polarimetry survey by George Wolf on the 24 inch Optical Craftsman Telescope at Mt. John Observatory in New Zealand during 1969.
- However, because of design and mechanical problems, that telescope was not delivered to New Zealand until 1970.
- This delay caused a change in venue to Kitt Peak National Observatory for a Northern Sky elliptical-polarimetry survey by Wolf during December 1969 and January 1970.
- The polarimeter was then moved to Flower and Cook Observatory in 1971.

#### The Mark I Polarimeter on Kitt Peak 0.9 meter



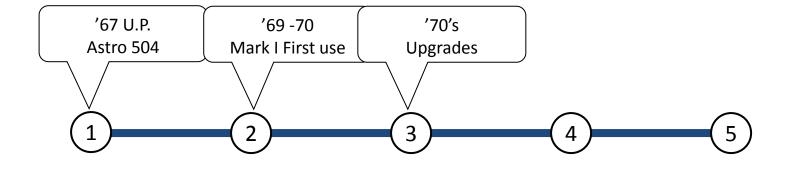
December 1969

## The Initial Mark I Polarization Survey

The initial elliptical-polarization survey (all four stokes parameters) at Kitt Peak of approximately 70 objects included:

- Magnetic A stars
- Polarized O and B stars
- Highly polarized stars of other spectral types
- Stars with peculiar spectra
- Intrinsic variable stars

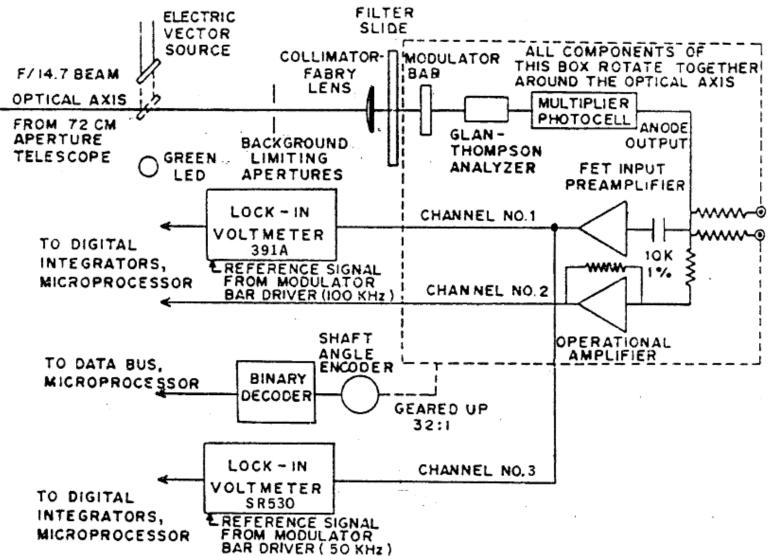
- Galactic and extragalactic objects with known synchrotron radiation
- Comet Tago-Sato-Kosaka
- Unpolarized standard stars
- Polarized standard stars



#### 3. 70's Polarimeter Upgrades & Redesigns

- 1972: Added a dual channel HP integrating digital voltmeter and digital clock.
- Closed circuit TV camera was trained on the digital displays of the clock and voltmeter.
- 1972-1973: An IBM key punch was interfaced with the digital voltmeter and clock.
- 1977? 1981 Polarimeter was redesigned. One photomultiplier tube, Photo Elastic modulator bar, internal polarization sources, Ithaco lock-in amplifier.

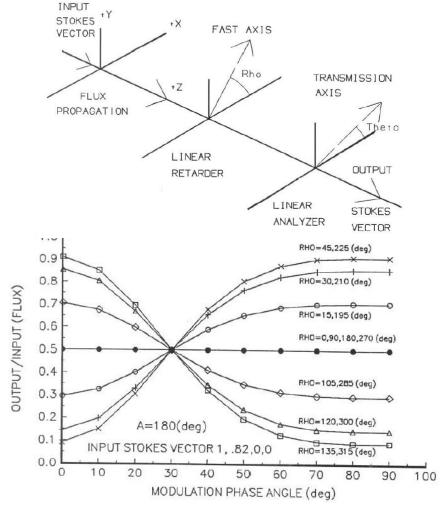
#### **Electro-Optical Components of the PEMP**



#### **PEMP Special Optical Component Details**

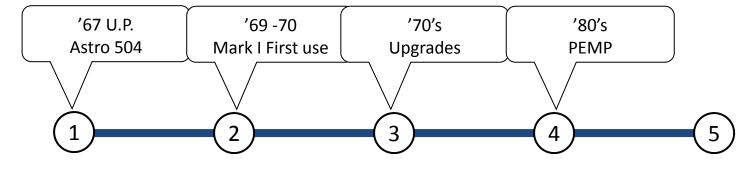
#### Hinds Photoelastic Variable Retarder & Controller





$$\begin{split} I_{d}(\Phi, t) &= \frac{1}{2}I_{T} + \frac{1}{2}P_{I}\cos(2\Phi)\cos(2\Psi) + \frac{1}{2}\sin(2\Psi) \bullet \\ &\{(P_{I}\sin(2\Phi)\cos\alpha + V\sin\alpha) \left[J_{o}(\beta) + 2\sum_{n=1}^{\infty}J_{2n}(\beta)\cos(2n\omega t)\right] \\ &+ (V\cos\alpha - P_{I}\sin(2\Phi)\sin\alpha) \left[2\sum_{n=0}^{\infty}J_{2n+1}(\beta)\sin((2n+1)\omega t)\right]\}, \end{split}$$

Blitzstein, et. al 1993



## 4. Later PEMP Operations and Improvements

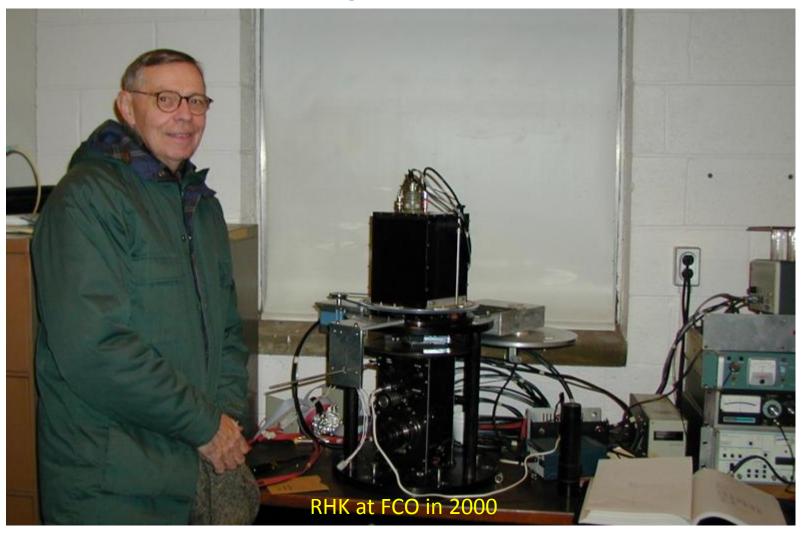
- 1981 Ohio Scientific microcomputer installed.
- 1981 Dave Bradstreet wrote first OSI polarization data reduction program.
- 1981 (August) PEMP Operations began
- 1982 Encoder added to Polarimeter to read and record azimuth.
- 1986 Ohio Scientific computer replaced with IBM AT and new interface.
- 1986 Stanford SR530 lock-in added for simultaneous Q-, U-, and V-Stokes vector operation.
- 1987 Temperature sensors installed into the Polarimeter and Lock-in.
- 1990's-2000 Remote operation completed: filter slide, PEMP orientation

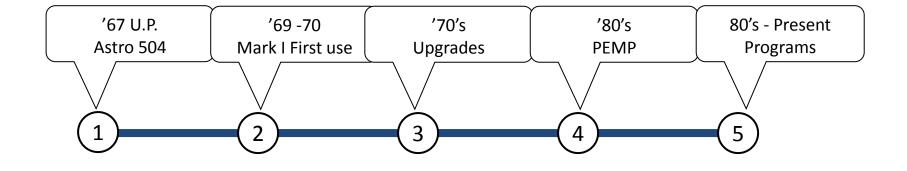
#### PEMP on 28-in. Cassegrain at FCO



Early 1980's

# **Remote Operation PEMP**



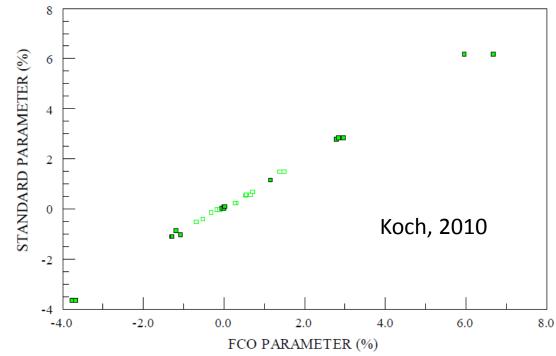


#### **5. 80's-Present Observational Programs**

Productivity: August, 1984 to October, 1992Close Binary Systems1673Luminous Late-Type Variables621Beta Cepheid Variables263Polarization Standards772Miscellaneous57Total Observations3386

Source: Blitzstein, et. al. 1993

# **PEMP Calibration**



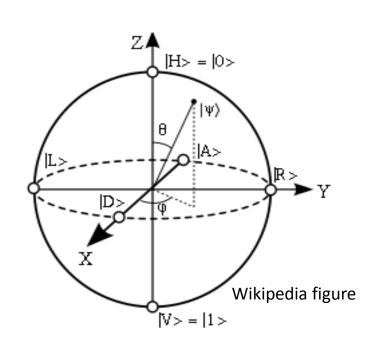
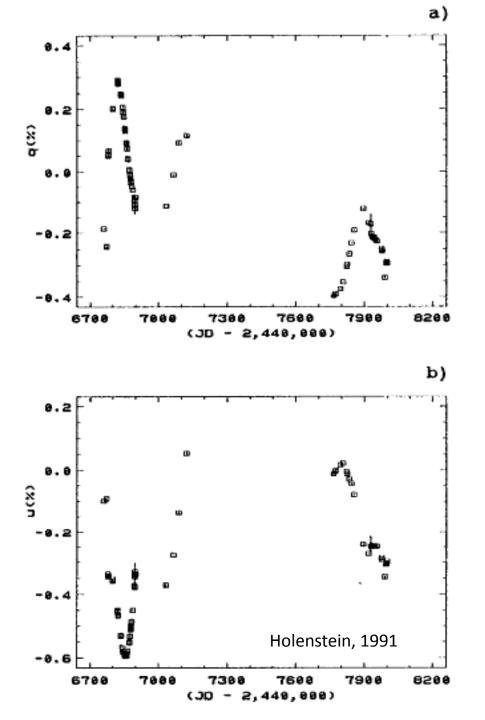


Fig. 52. The calibration of the FCO values of the green Q and U polarization parameters over the interval, July 1983 through May 1984. Filled symbols refer to Q and open ones to U. Linear regressions lead to: Q(std) = +0.0301 + 0.9668Q(FCO), U(std) = +0.0176 + 0.9293U(FCO).

Null and non-null standard stars for linear calibration

Null standards and rotation of the Poincaré sphere for circular calibration



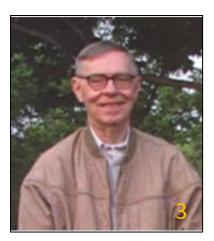
# Luminous Late-Type Variables

- Alpha Ori (R-bandpass shown)
- Others: 119 CE Tau,
  6 BU Gem, Ψ<sup>1</sup> Aur,
  72 Leo, V CVn, α Sco,
  α<sup>1</sup> Her, μ Cep, VV
  Cep, β Peg

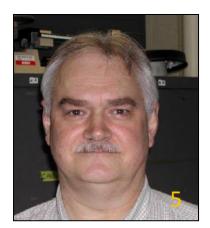
## Other key Penn Polarimeter Contributors

- 1. Bill Barrie
- 2. Dave Bradstreet
- 3. Robert H. Koch
- 4. Robert E. Smith
- 5. Richard J. Mitchell









# Summary

- Penn Polarimeters evolved over 30 years and made use of the best available technology during that time.
- Product of a collaboration of many Penn astronomers and staff from 1967 to 2000.
- Elliptical polarimetry, using a photoelastic modulated bar, was a productive activity at Flower and Cook Observatory, a location near a major US city.
- Objects all over the HR diagram down to about 11<sup>th</sup> magnitude were studied polarimetrically.
- The precision in latter years of better than <u>+</u> 0.01% for a 7<sup>th</sup> magnitude star.